Field-of-View Calibration of the Microwave Limb Sounder on the Upper Atmosphere Research Satellite

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Abstract: This paper describes the field-of-view (FOV) calibration of the Microwave Limb Sounder (MLS) on board NASA's Upper Atmosphere Research Satellite (UARS). Calibration data are derived from measurements and analytical models, combined with in-flight data. Particular emphasis is given to pointing calibration and estimation of faiside lobe levels and radiance offsets, using both the Moon and tile residual signals when the FOV is pointed high above Earth's atmosphere.

INTRODUCTION

The MLS represents the first implementation of microwave limb sounding from space (Barath, 1993). The MLS senses—thermal emission in 6 bands with passive radiometers at 63, 183 and 205 GHz. Its three-mirror antenna system is a 1.6 x 0.8 meter offset Cassegrain matched to diffraction-limited optics and having a surface accuracy of 25 microns. Knowledge of the antenna patternshape and relative pointing offsets between radiometers is critical for accurate retrievals of temperature, pressure and minor constituent profiles.

A companion presentation (Jarnot,1994) describes the instrument, its operation, and spectral and radiometric calibration to date, for which the MLS data validation program has indicated absolute accuracy of 5% or better, Complete details of all pre-launch calibrations are in the MLS (calibration Report (Jarnot, 1991).

FIELD-OF-VIEW (FOV) CALIBRATION

FOV calibration consists of the determination of the response of the MLS, as a function of direction, to received power. Table 1 summarizes the antenna subsystem performance, and shows the designations of radiometers (R1-3) and bands (B-6). Dependences of the FOV on scan angle, orbital conditions, and signal frequency within bands, expected to be small compared to its angular dependence, were also characterized for further corrections in the data processing.

The 4π solid angle domain of the FOV functions was divided into two parts based on the rates of change of both the FOV and the atmospheric signal, and on whether we characterize the FOV by model or by measurements. These conform to the partitioning of flight data processing into two "levels":

Level 2: In the solid angle Ω_A , extending ± 0.1 radian $\pm 5^\circ$ about the direction to the nominal center of the scan (limb tangent height $h_T=30\,$ km), both the atmospheric signal and the FOV vary rapidly with angle. FOV calibration] in this region is performed at Level 2 by a forward radiance model whose inputs include FOV response vs angle from boresight, for each radiometer.

Level 1: Outside Ω_A , the FOV response is $< 10^{\circ\circ}$ of its peak value; however, the beam solid angle is so much less than 4π that the average FOV level must be characterized to $10^{-7.5}$. As this was only possible for Band 1 with the available test equipment and far-field range,

many of the calibration data depend on analytical models. Since much of this portion of the FOV views ^{1a}-diances which are Illatively constant and/or difficult to model (c.g. Earth, MLS/UARS structures, and spat.c), calibration is supplied to Level 1 as quasi-constant transmissions and radiance offsets.

LEVEL 1 ANTENNATRANSMISSION AND RADIANCE OFFSET

The radiance from the antenna which is incident on the radiometer limb port, for channel in radiometer r, is given by:

$$\begin{split} \hat{I}_{i}^{A} &= \rho_{\tau}^{P} \rho_{r}^{S} \rho_{\tau}^{T} \eta_{\tau}^{AA} \eta_{\tau}^{P} \hat{I}_{i}^{I} \\ &+ \rho_{\tau}^{P} \rho_{r}^{S} \rho_{\tau}^{T} (1 - \eta_{\tau}^{AA}) \eta_{\tau}^{P} \hat{I}_{r}^{SL} + (1 - \rho_{\tau}^{P}) \rho_{r}^{S} \rho_{\tau}^{T} \eta_{\tau}^{P} \hat{I}_{\tau}^{OP} \\ &+ \rho_{\tau}^{S} \rho_{\tau}^{T} (\eta_{r}^{S} - \eta_{\tau}^{P}) \hat{I}_{r}^{SP} + (1 - \rho_{\tau}^{S}) \rho_{\tau}^{T} \eta_{\tau}^{S} \hat{I}_{\tau}^{OS} \\ &+ \rho_{\tau}^{T} (\eta_{\tau}^{T} - \eta_{\tau}^{S}) \hat{I}_{\tau}^{SS} + (1 - \rho_{\tau}^{T}) \eta_{\tau}^{T} \hat{I}_{\tau}^{OT} \\ &-1 (1 - \eta_{\tau}^{T}) \hat{I}_{r}^{ST} \end{split} \tag{1}$$

where

 k_{τ}^{k} = Reflectivity of reflector k = P, S, T(Primary, Secondary, Tertiary)

 η_r^{AA} : Beam efficiency of the antenna system: the product of scattering (η_r^{AS}) and diffraction (η_r^{AD}) from the primary aperture plane

 η_{τ}^{Sk} = Spillover efficiency of reflector k with measured feed patter n r

 $\dot{I}_{i}^{L} = \text{Limb Radiance, channel 7}$

 $\dot{I}_{\bullet}^{SL} = -$ Radiance from outside FOV measurement angle Ω_A , in the limb hemisphere

 \dot{I}_r^{Sk} . Radiance illuminating spillover solid angle for reflector k

 I_c^{Ok} Temperature of reflector k

Equation 1 can be derived by projecting all apertures to the limb portplane, noting through which reflectors and into which spill over each solid angle terminates.

A simplified form of this equation is inverted in Level I calculation of calibrated radiances:

$$\dot{I}_{i}^{L} = \frac{1}{\eta^{A} \rho_{\tau}^{A}} \left(\dot{I}_{i}^{A} - (1 - \rho_{\tau}^{A}) \dot{I}_{r}^{OA} - (1 - \eta_{\tau}^{A}) \rho_{\tau}^{A} \dot{I}_{\tau}^{SA} \right), \qquad (2)$$

For this simplification, reflectivities $\rho_r^{P,S,T}$ are combined into a single-ohmic transmission ρ_r^A , and internal s})illover. losses similarly with scattering and diffraction losses into the antenna transmission η_R^A . Likewise, the radiances aucf physical temperatures in equation 1 are combined into effective radiances \hat{l}_r^{OA} and \hat{l}_r^{SA} .

Transmissions and radiances were combined over all loss mechanisms and antenna components independently for each radiometer. The calculation of $\eta_r^A, \rho_r^A, \hat{J}_r^{OA}$ and \hat{J}_r^{SA} from the measured ρ_r^k and η_r^{SA} , modelled η_r^{AD} , and for reasonable pre-launch estimates of \hat{J}_r^{SL} , \hat{J}_r^{SR} and \hat{J}_r^{OK} is discussed in (Jarnot, 1991)

Table 1: MLSAntennaFOVPerformance from Ground Calibration and Analytical Models. Knowledge values are 30

v	O		
Parameter	R1,B1 63 GHz	R2,B2·4 205 GHz	R3,B5 6 183 GHz
Half-Power Beam Width (HPBW) (vert.) / °	0.206	0.064	0.077
K n owledge	0.002	0.001	0.001
HPBW (1101.)/°	0.43	0.145	0.152
Knowledge	0.008	0.003	0.002
BeamEfficiency	0.91	0.9	0.91
Knowledge	0.01		
Polarization(angle between of E and vertical at 30 km tangent point) / °	I 4.	91.	2
Peak cross-polarization / dB FOV direction (dFOV) knowledge (Wt.) / °	-30	- 19	-20
absolute,B1 to optical reference cube	0.0036		
relative to B1		0.0016	0.0021

Contributors to Level 1 FOV Bud. t

Four mechanisms contribute to antenna transmission and radiance offsets:

- Ohmic loss within the antenna reflectors, which was inferred from reflectivity measurements in allbands. Using calculated reflectivities of a silver-plate standard, the wo [st-case reflectivity (205 GHz, Primary Reflector)] was $R_{1,R2} = 0.9956$ with standard deviation of 0.0008, Combining this error with the systematic uncertainty in the calculation of absolute reflectivity, and the time-k, aryill, g-reflector temperatures, gives ohmic efficiencies ρ_r^A ranging from 0.99230.003 in B1 to 0.98930.004 in B2-4.
- 2 Spillover at antenna reflectors and radiometer apertures. This was obtained by integrating measured feed patterns to the three port edges and to the projected outlines of the antenna reflectors. Uncertainty in normalizing power dominates errors in the limb port baffle transmissions appearing in radiometeric calibration, but tends to cancel in the spill over contribution to η_{τ}^{A}
- 3 Scattering by antenna surface irregular ities, which was estimated using the contour measurements made during the manufacture of the reflectors, on both coarse (2-D) and fine (1-D) grids. The first dataset was used to calculate errors of form. Fourier analysis of tile second provided correlation length and rms deviation (Marx, 1990) for estimating, mathe Ruze statistical model (Ruze, 1966), the scattered power which was hidden from FOV patterns with noise floor encountered on the far-field range. This fraction of the scattered power was budgeted as antenna transmission. Variations in transmit terpower and at mospheric attenuation on the fill-field range make this fraction not monotonic in frequency.

Mean values of the first 3 or 4 sidelobes measured for B2 4 matched the predicted scattered pattern envelope. Model errors of 10% uncertainty in the rms amplitude c and 15% in the correlation length L give 40% uncertainty infractional

Table 2: Antenna Transmission and Radiation Offsets for each MLS Radiometer

$\begin{array}{c} Contributor \\ \rho_r^A \text{(reflectivity)} \\ \text{scattering} \\ \text{edge diffraction} \\ \text{Primary spillover} \\ \eta_r^A \end{array}$	B1 0.9923 0.999 0.977 0.954 0.931		B5- B6 0.992 0.988 0.992 0.940 0.921
is A i o A range in orbit) Ohmic Diffraction/Scattering Total Offset	252.3 K	243.1 K 256.2 K 203,313] H 2.8 K 5.8 K 8.6 K	251.3 K

scattered p ower 1 - $\eta_\tau^{AS},$ corresponding to 0.7 K in radiance offset for the worst case, B5-6. "

4 Edge diffraction, for which a model was developed by applying the Geometrical Theory of Diffraction (GTD) to the Primary Reflector leading edge (the limiting aperture at -15 dB taper). GTD patter 0s match the FOV shape predicted by the aperture-field method and seen in the n reasured FOVs at O. 6° for Band 1, where spillover is greatest and this diffraction is most pronounced. Edge diffraction efficiencies η;") appearing in equation 1 ranged from 0.9773 0.009 in B1 to 0.9984 0.001 in 112-4.

Table 2 summarizes Level 1 FOV parameters from ground calibration.

FOV CHARACTERIZATION

FOV pattern measurements were performed using frequency locked transmitters as the signal sources, and were digitally recorded on ground support equipment computers in spherical polar coordinates. Feed patterns were measured in both subassembly and radiometer configurations to verify proper antenna illumination, alignment sensitivity and aperture spillover levels.

Secondary FOV Measurement

FOV patterns of the MLS sensor were characterized at 10 scan angles and - 5 frequencies within each band. Patterns were measured at cuts spaced in azimuth by 22.5°, but at closer—spacing near features of special interest and to verify predictions of the analytical Hodels. Polar angle resolution varied from O.O1° in the far sidelobes to HPBW/70 on the main lobe. The ensuing random errors combined with systematic due to range equipment, transmitted power drift and varying temperature fields and gravity loads on the MLS, to give < 3.5% errors in HPBW.

Boresight directions (d FOV) at 63 and 205 GHz were measured to 15 arcsec accuracy, relative to a sensor alignment cube, using atheodolite in conjunction with the RF patterns. A more stringent ? at csec knowledge of relative dFOV coincidence between radiometers was measured with near-simultaneous patterns using two transmitters.

Figure 1 shows a measured pattern of Band 4 in the limb vertical plane. Noise floors for the 2 polarizations represent the 3km and 1km far-field tanges used. Since the limb radiance variation is significant only in the vertical direction over the ±6° domain treated by Level 2, measured FOVS were collapsed, i.e. integrated along

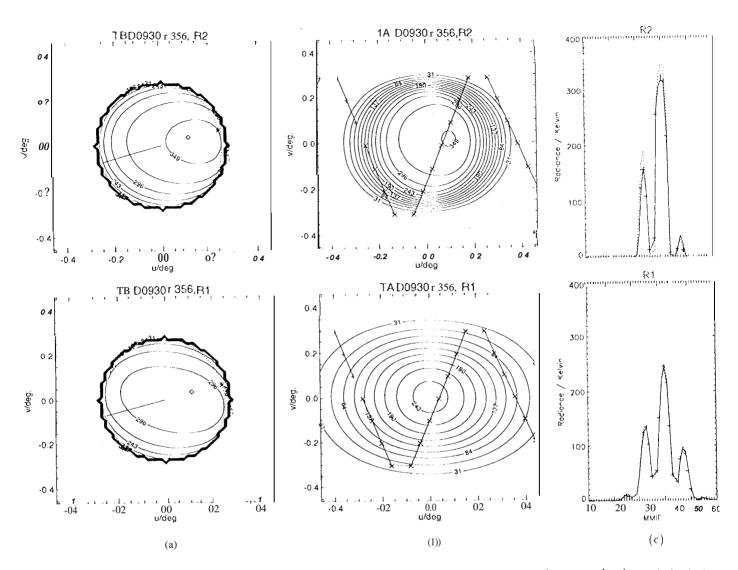


Figure 2: Example of FOV scan through Moon: (a) Model radiance incident on far-field sphere of antenna, showing polarization) dependence and model angular resolution: (b) Convolved limb radiance with scan pattern: (c) Radiance time series before and after iterative solution for dFOV and model gain

attributed to incomplete characterization of the R3 feed patterns and to launch shift. After $2^{1}/_{2}$ years of operation, no instabilities or drifts have been traced to FOV calibration.

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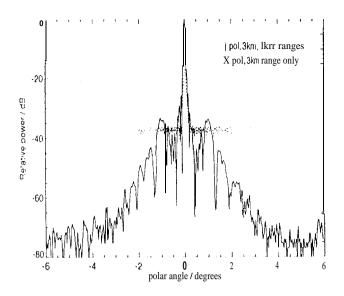


Figure 1: Band 4 Limb Vertical FOV pattern

the horizontal direction, to provide 1-dimensional FOV functions for the forward model.

Dependences of FOV on scan angle and frequency within the band at e consistent with known measurement errors and within the requirements of flight data processing algorithms.

Forthelarge (> 100 I{) signals in Bands 1 and 4-6, variations with scanangle and IF meet the 0.5% functional requirement when convolved with canonical atmospheric radiance kernels at band centers, and compared as differential antenna radiances, ΔT_A . For the CIO signal (Bands 2 and 3), the scatter in the ΔT_A curves is interpreted as the peak to peak excursion of a random error whose standard deviation $\sigma \sim (1 \sim \text{eak-to-peak})/\text{fi} < 0.05 \text{ K}$.

REFINEMENT OF FOV CA LIBRATION FROM IN-ORBIT DATA

Absolute Knowledge of dFOV

soou after launch, MLS pointing was validated at Level 2 by comparing measured radiance spectra growths vs. tangent hri.gilt, inferred from dead-reckoned pointing, to the forward model. A large offset of 0.12", attributed to launch shift or placement uncertainty between MLS and UARS alignment cubes, was found and corrected in the calibration data

lligh-Altitude Radiances and 5-20° Roll Maneuvers

Another early refinement to pre-launch FOV calibration was adjustment of scene radiances in equation 1 to make band-averaged radiances at the Highest-altitude minor frame (MMIF) match the Planck function values for space radiance. Most of this adjustment occurred in the assumed brightness temperature of UARS seen by MLS Primary spill over. Reductions in total offset radiance ranged from 3.2 to 5.7 K. This was validated by radiance data during special scans to the antenna home position ($h_T > 250\,$ km) and during ~ 1 orbit of observation with UARS rolled 5° and 20° above the nominal attitude.

Comparison of 1{2 and R3 Retrieved 0₃ Profiles

Ozone profiles retrieved using pre-launch calibration data for R3 exceeded those for R2 by as much as 15%. Both sideband ratio and antenna transmission were adjusted to reduce this discrepancy to < 5%. For the latter, the pre-launch values of R3 baffle

Table 3: Coincidence of Vertical dFOV between Radiometers

	R2-R1 / °		R3- R2 / °	
	$d\epsilon$	3σ	$d\epsilon$	30
Pre-Launch	-1 0.006	0.0016	-0.007	0.0026
forom Moon scans	-10008	0.009	-10.004	0.007

transmissions, and the concomitant reflector spillover losses, were replaced by those of R2. Also, an overly conservative estimate of -42dB for the noise floor of FOV patterns was replaced by -50dB. I.E., more information lay in the pre-launch FOV than had been previously budgeted; hence η_{183}^{AA} could be increased accordingly. Further steps to reconcile R2 and R3 ozone treated deviations from pre-launch values of dFOV coincidence

Estimation of dFOV Coincidence from Moon in FOV

The envelope of FOV over the avail able MLS scan range intercepts the Moon approximately 17 clays per year. On four days the MLS scan pattern was altered for ~1/4 of each of 612 orbits per day, to let the Moon drift through the FOV at controlled scan angles well above the atmosphere. Calibrated radiances were compared to radiances from a lunar microwave model (Keihm, 1983) and convolved with the measured FOVs and with a step function which models smear'illg in azimuth due to the integration time. One can relate measured radiances to the model map and its gradient by two pointing angles and a model scaling factor (this sealing error mimics uncertainty in the 10V gain, but is bounded by the uncertainty in high-altitude radiances). These 3 parameters are estimated by minimizing the squared residual sum for the 30 or so MMIFs in each Moon crossing.

Figure 2 shows model radiance maps before and after convolution. The special scan pattern shown was tailored to minimize the a priorivariance of retrieved vertical pointing, given uncertainty in UARS ephemeris predictions and the unavoidable horizontal drift due to orbital motion. The rightmost panels superimpose measured radiances ou model values before and after the retrieval, showing the high signal-to-noise ratio of this technique.

Line 2 of Table 3 results from pooling dFOV between radiometers over all scans to date. Some systematic errors canceling the subtraction of retrieved angles to obtain the dFOV between Rlandthe other two radiometers. Uncertainties for the insorbit method currently remain several times those claimed for the ground calibration. The change in dFOV between R2 and R3 is statistically significant and agrees to 0.003° with the offset required to account for the disc repancy between R2 and R3 ozone retrievals. This change is attributed to launch shift and has been used to update data processing software. The scale factor retrieved for 63 GHz differs from 1 by an amount consistent with the Moon model's accuracy at 90 GHz (Keihm, 1983), but the larger deviations for R2 and R3 may suggest some residual error in far sidelobe level of the FOV functions. This and characterization of model uncertainties are 00W under investigation.

CONC1,[JS10NS

All FOV characterization eventually required by the software for flight data processing was identified early in the instrument's development, and nearly all was provided by ground calibration. Exceptions were absolute pointing error, due to launch shift between MLS and UARS references, and radiance offset, due to incomplete knowledge of background stray radiance. The latter is spectrally flat and therefore negligible in Level 2 processing. Subsequent in orbitrefinements have addressed 1{1-1{3} ozone differences, and are